

#### LA-UR-21-28486

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Title: Wave Generation and Wave-Particle Interaction Using Space-Based,

RF, Linear Electron Accelerators

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Intended for: The twenty-second edition of the International Conference on

Electromagnetics in Advanced Applications (ICEAA 2021), 2021-08-08

(Honolulu, Hawaii, United States)

Issued: 2021-08-25



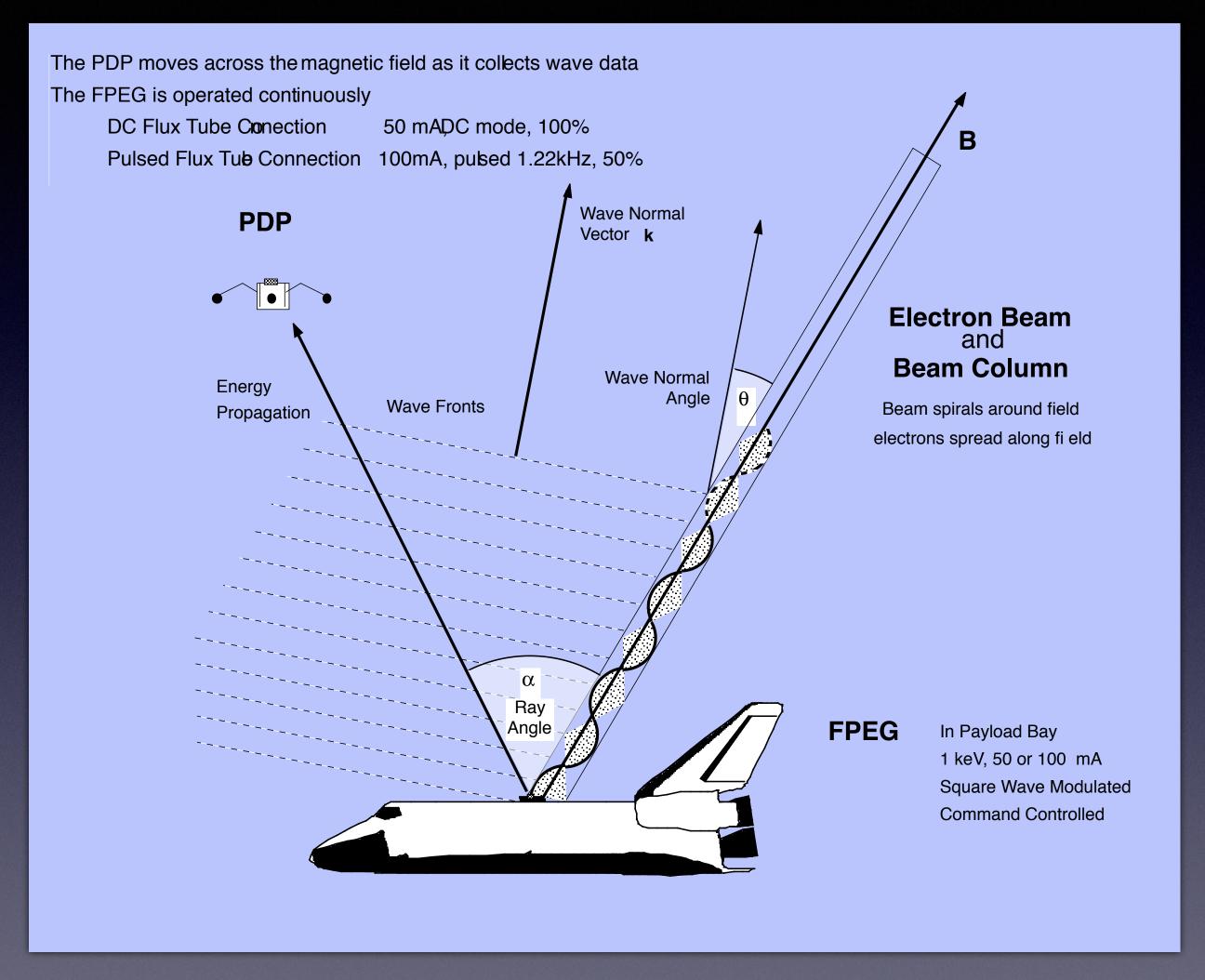


## Wave Generation and Wave-Particle Interaction Using Space-Based, RF, Linear Electron Accelerators

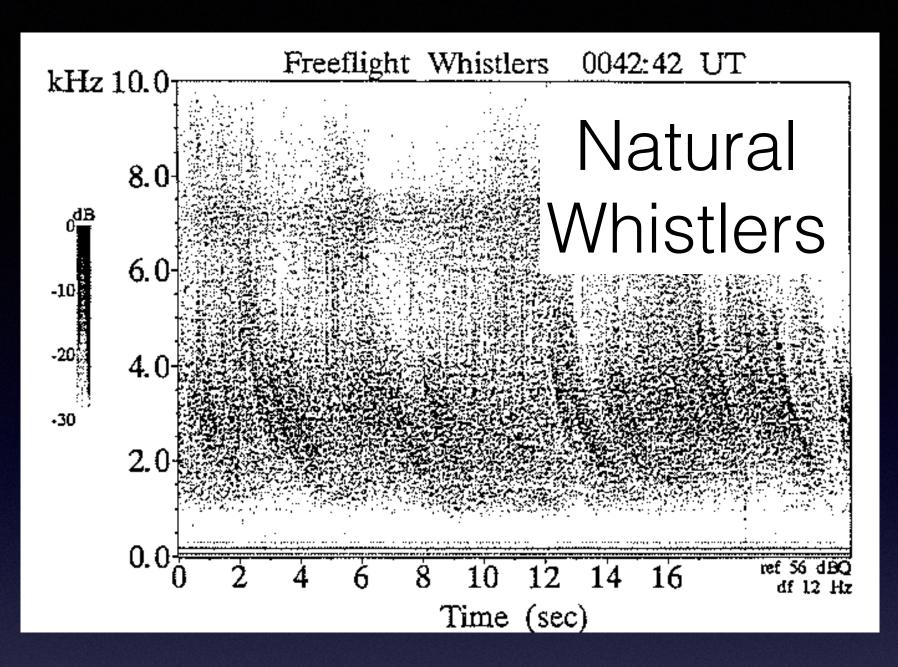
Geoffrey Reeves and the Beam PIE and RBR-ACES Teams

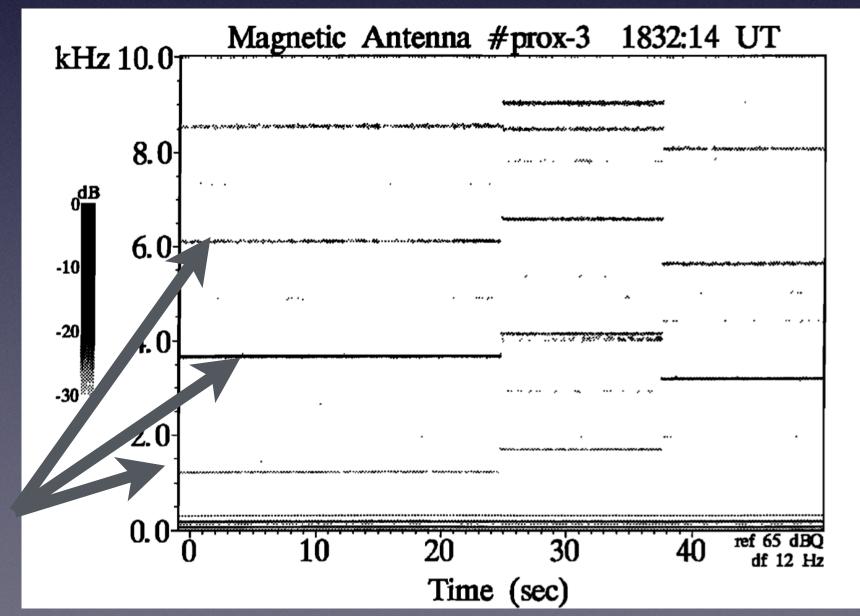
- Active experiments in wave-particle interactions using electron beams date back to the 1980s but went through a prolonged period of inactivity
- Van Allen Probes and other missions have vastly increased our understanding of wave particle interactions and their importance
- Technologies for beam generation, wave receivers, and particle detectors have seen huge advances in capability
- There's an opportunity for a new generation of active experiments to produce fundamental discoveries under controlled conditions not available in the natural environment

### STS-3 and Spacelab 2



Stimulated Emissions & harmonics



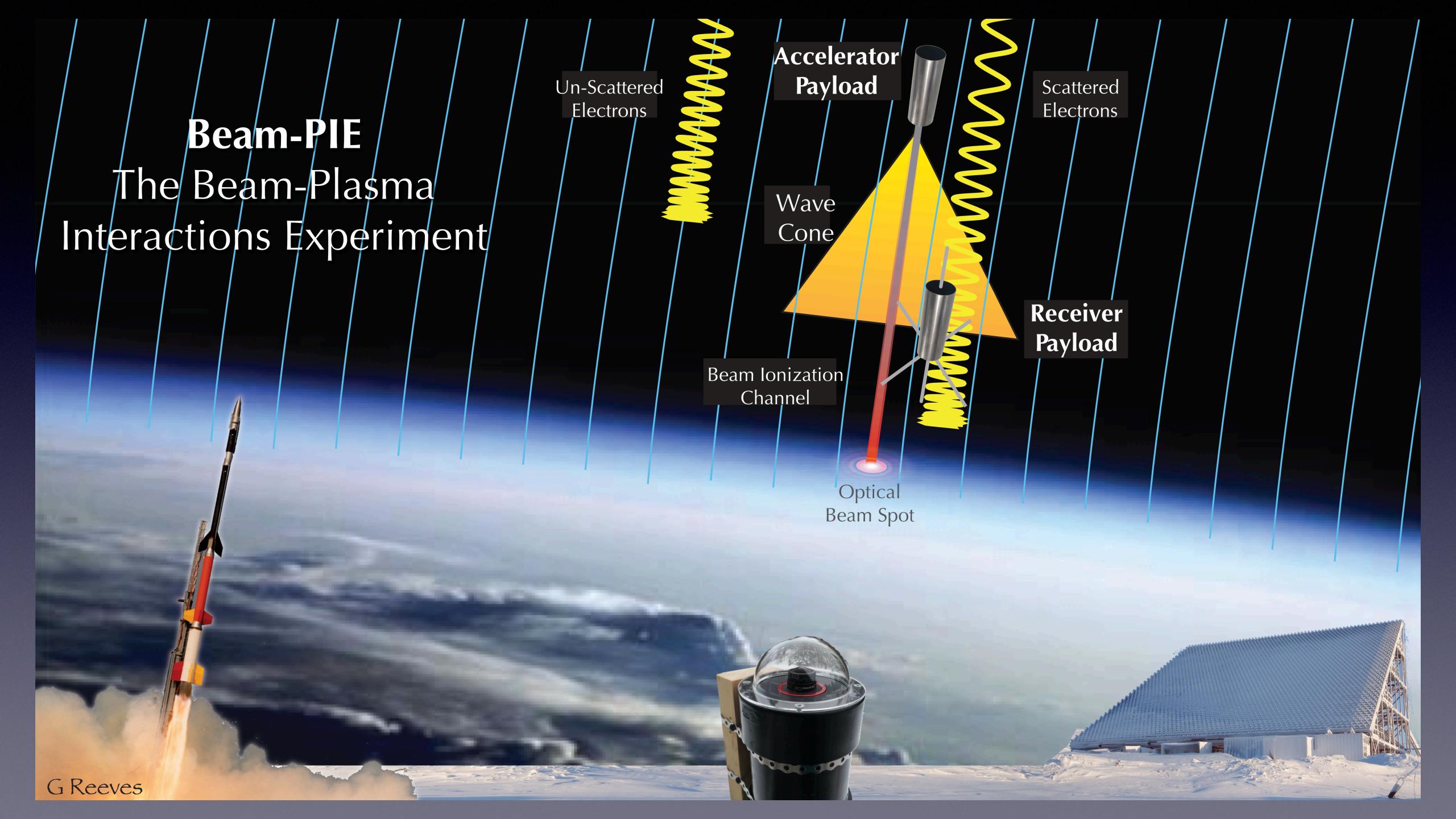


## Shuttle and rocket experiments confirmed:

- waves are stimulated by coherent Cherenkov emission
- observed waves are consistent with whistler mode
- spatial coherence and harmonic structure showed the beam propagated a long way without major instabilities (contrary to simulations)
- dependence on pitch angle, frequency, duty cycle, etc. were qualitatively consistent with theory [Harker and Banks, 1986]

## They could not:

- vary the energy of the beam
- measure waveforms for wave normal angle, ellipticity, planarity, etc.
- resolve spectral width of the emissions (important for pitch angle diffusion and non-linear scattering calculations)
- measure the effect of the waves on particles
- measure frequencies above ω<sub>ce</sub> where X-mode waves are predicted

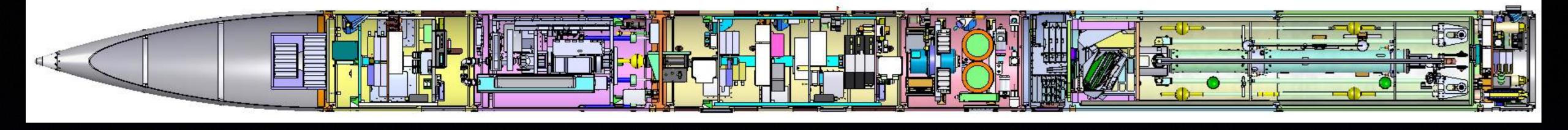


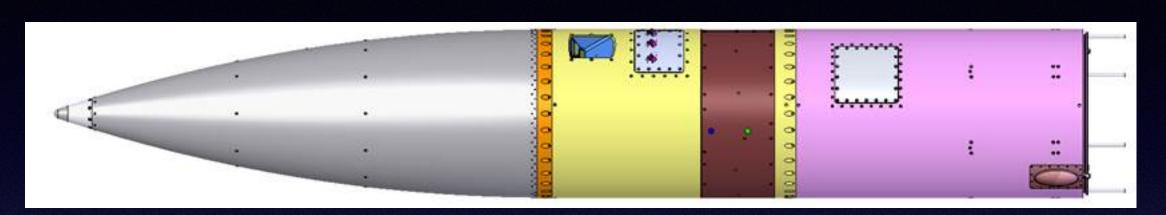
### Beam PIE Basics

- Joint LANL and GSFC collaboration
- We use a pulsed electron beam to generate waves using various frequencies, duty cycles, and energies
- A second payload measures waves, particles, and background plasma parameters (B, density, temperature)
- Measurements are capable of <u>quantitative</u> comparison against linear theory and simulations
- Non-linear effects can be resolved
- Nominal Launch is March 1, 2022

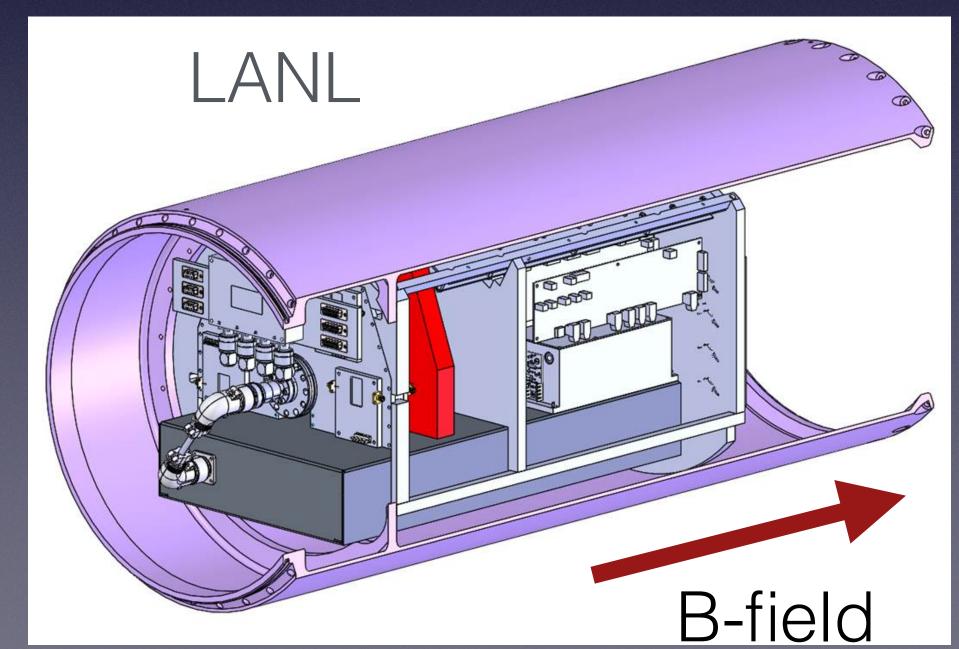
## Beam PIE Components

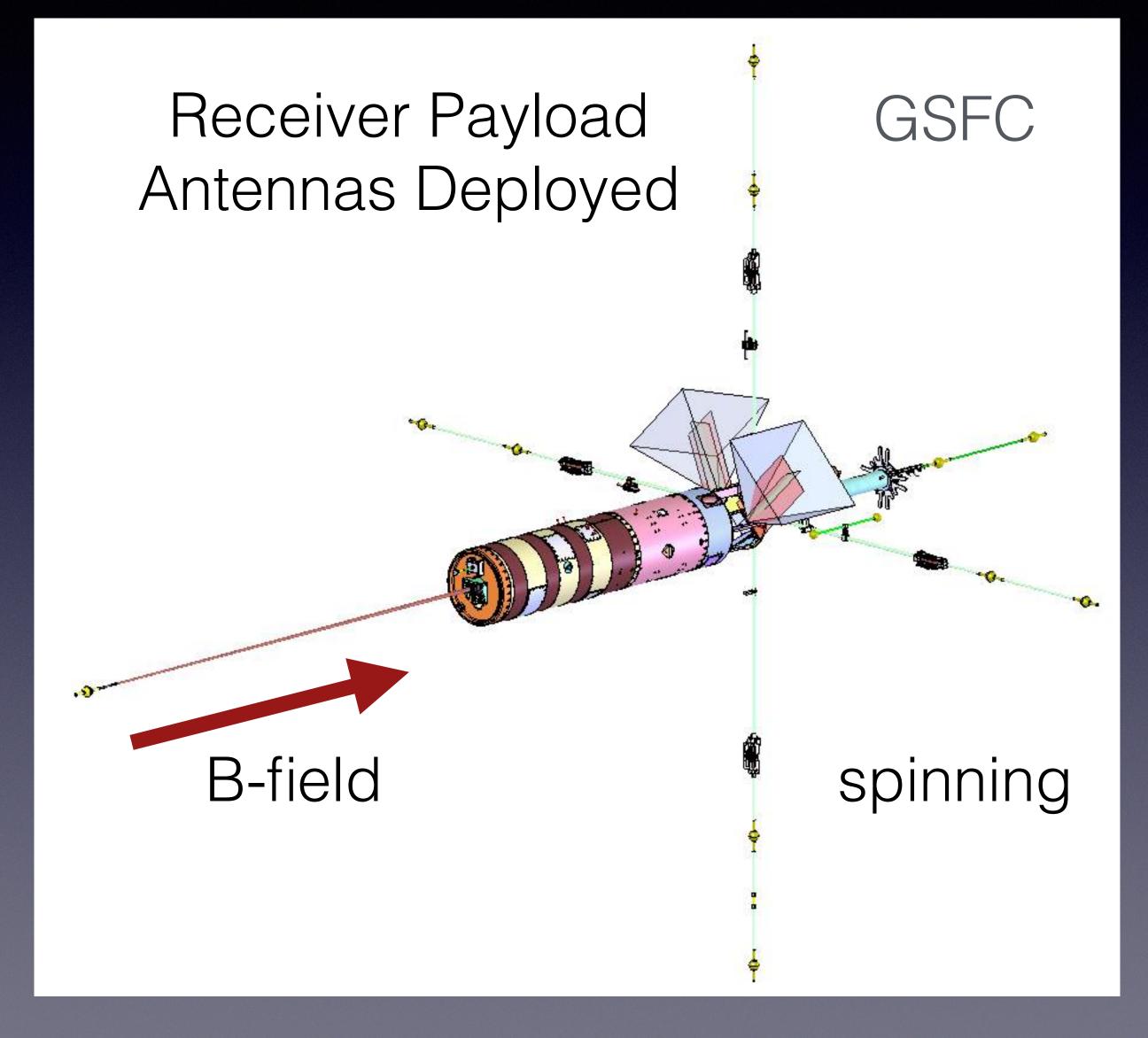
- Accelerator Payload: no measurements, spacecraft charging mitigated by maximizing surface area (e.g. nosecone)
- Wave Receiver Payload: full E and B waveform capture at VLF frequencies, 1D E component at HF frequencies
- Effect of Waves on Particles: two detectors to look for electron pitch angle scattering
- Ground Segment: will measure optical emissions from beam interaction with the atmosphere to assess beam propagation
- Theory & Simulation of: Accelerator, Beam Propagation, Wave Generation, & pitch angle scattering

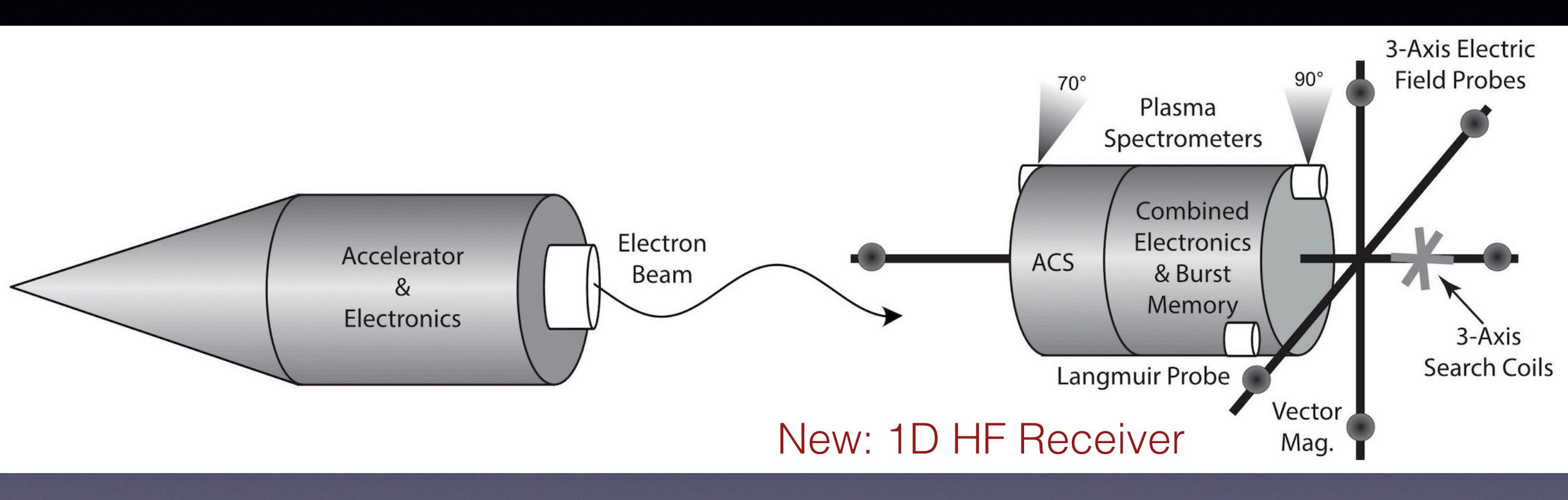


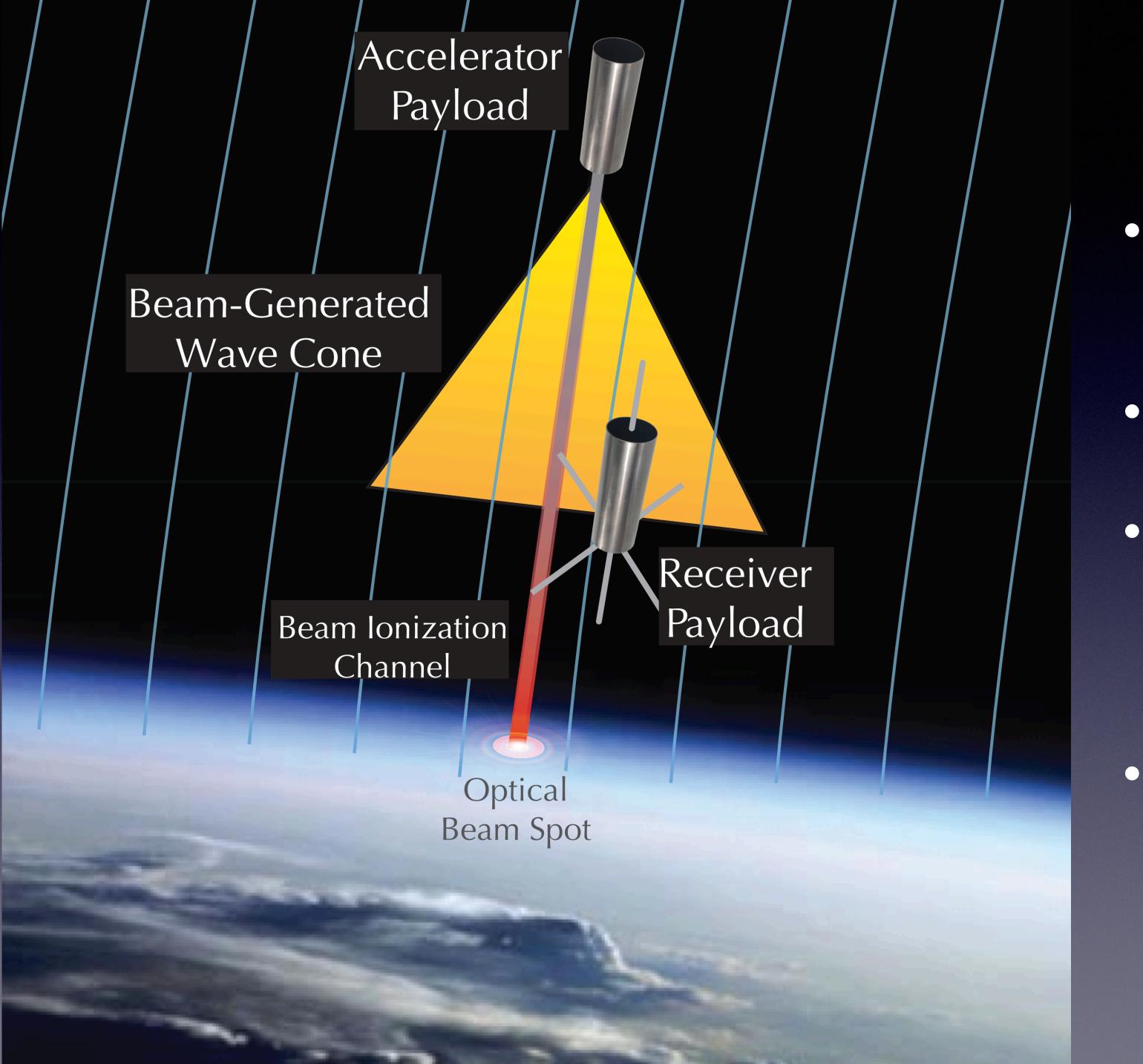


Accelerator with Nosecone attached to minimize charging





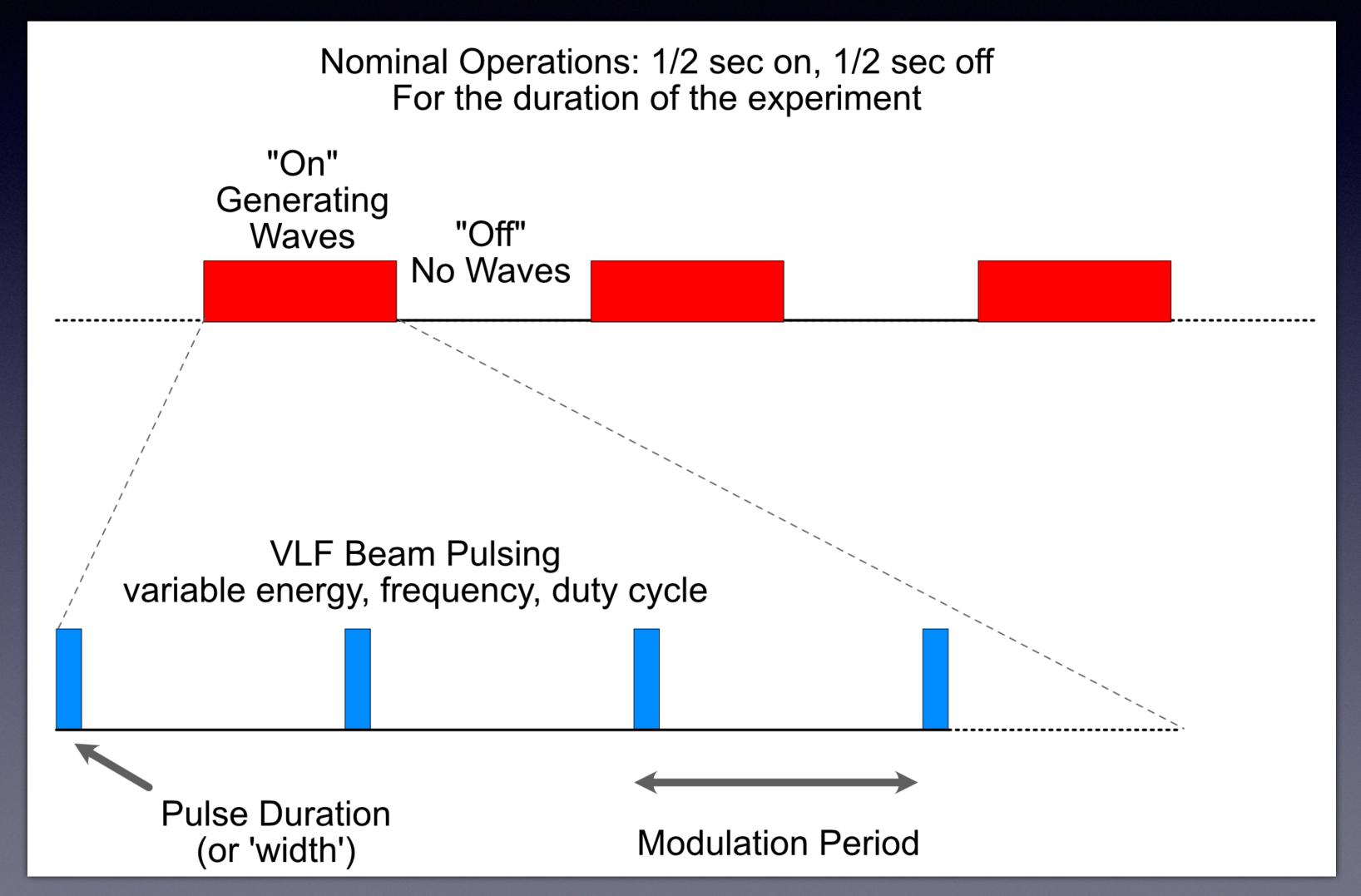




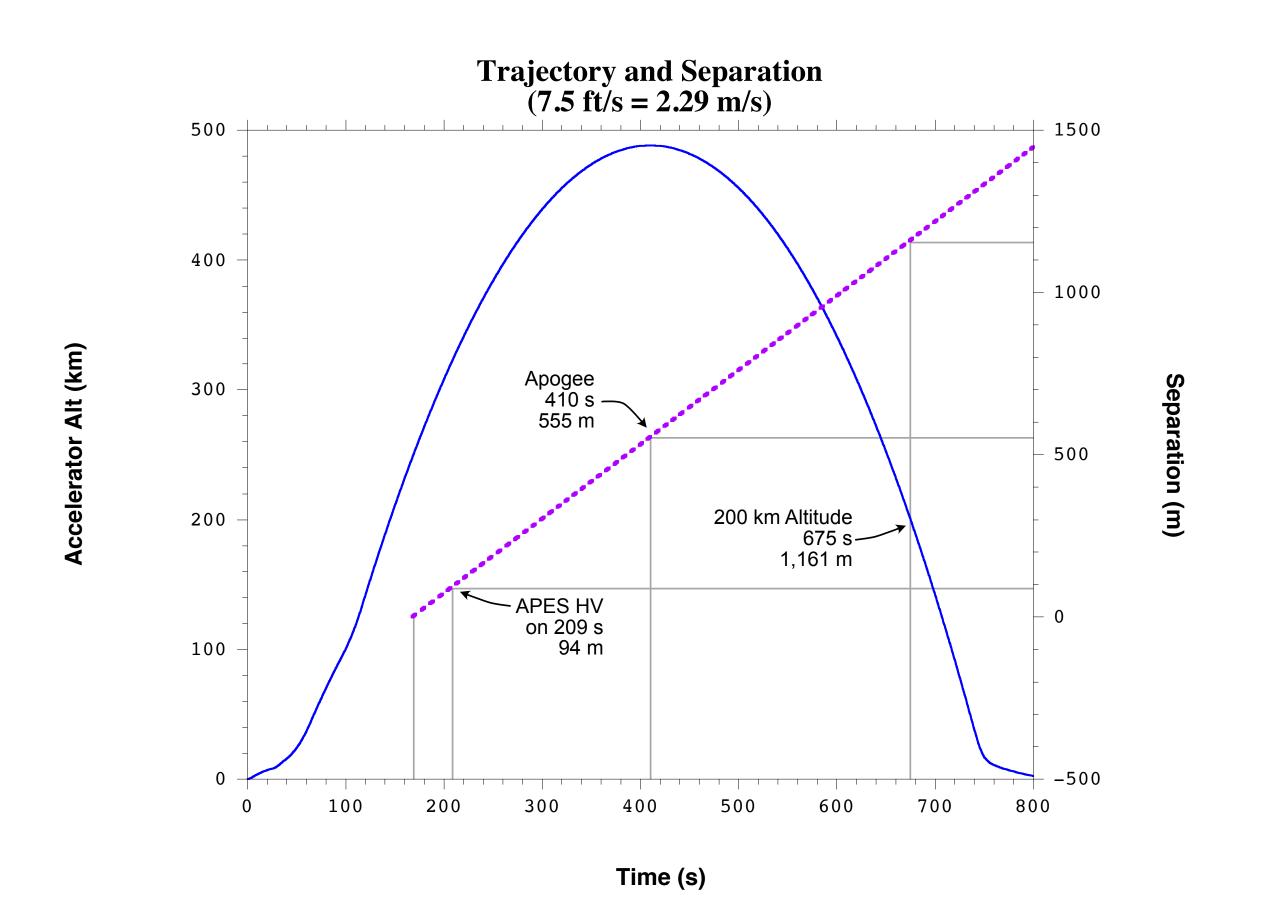
- Accelerator and receiver aligned along B
- Gradually separate
- We are not trying to measure the beam with the receiver
- We measure the waves in a cone around B

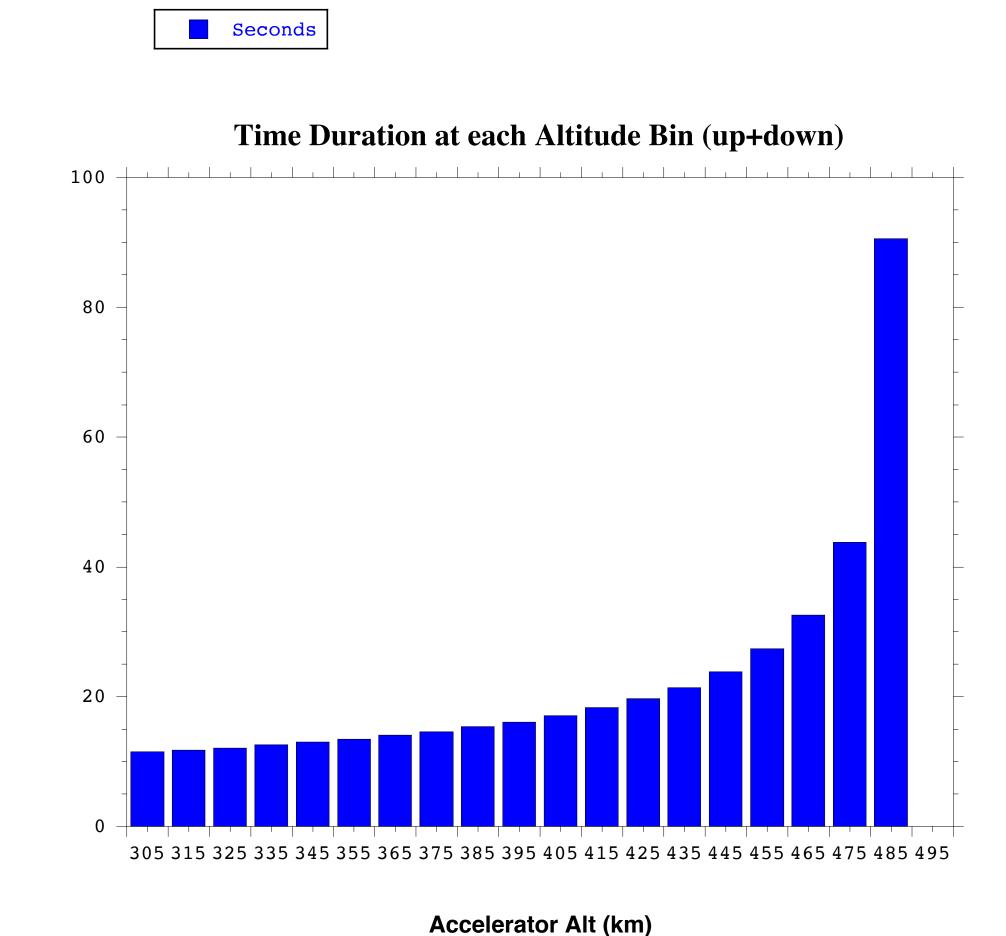
#### The beam pulsing scheme lets us

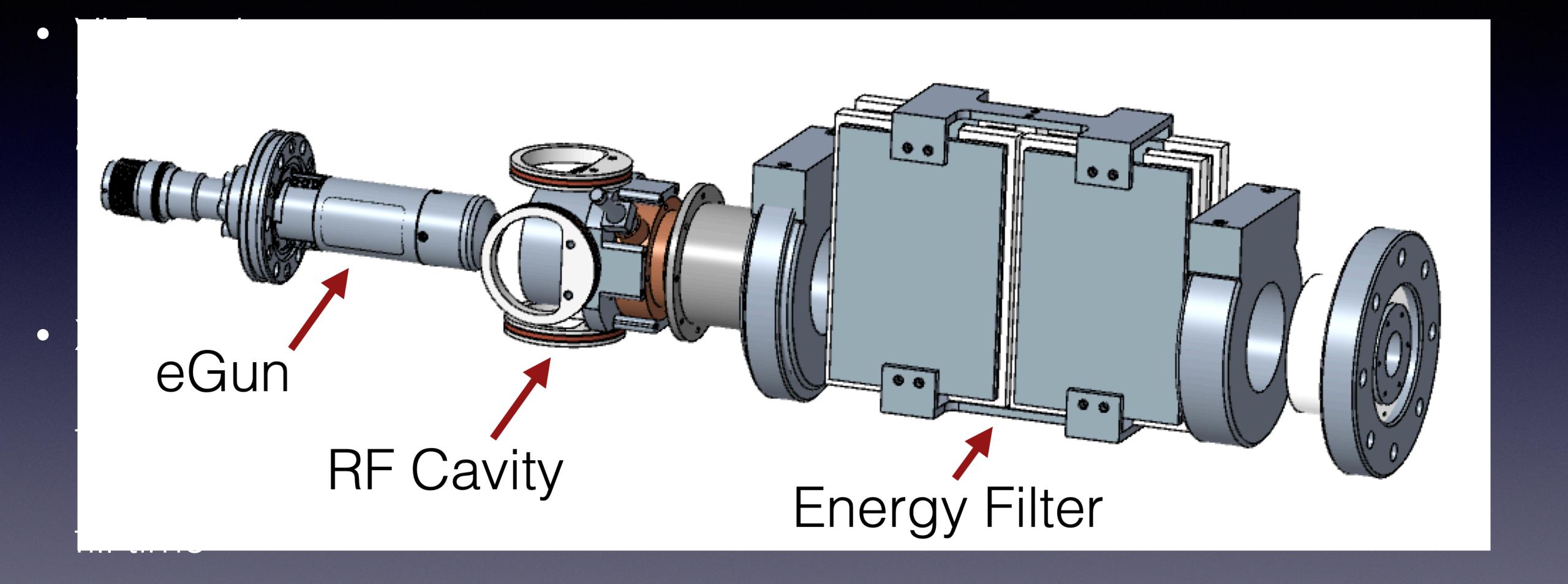
- vary energy, frequency, and duty cycle independently
- compare wave & particle measurements when the beam is on with background measurements when the beam is off



# Each ½ second (energy, frequency, duty cycle) sequence will be repeated at different separations and background plasma conditions

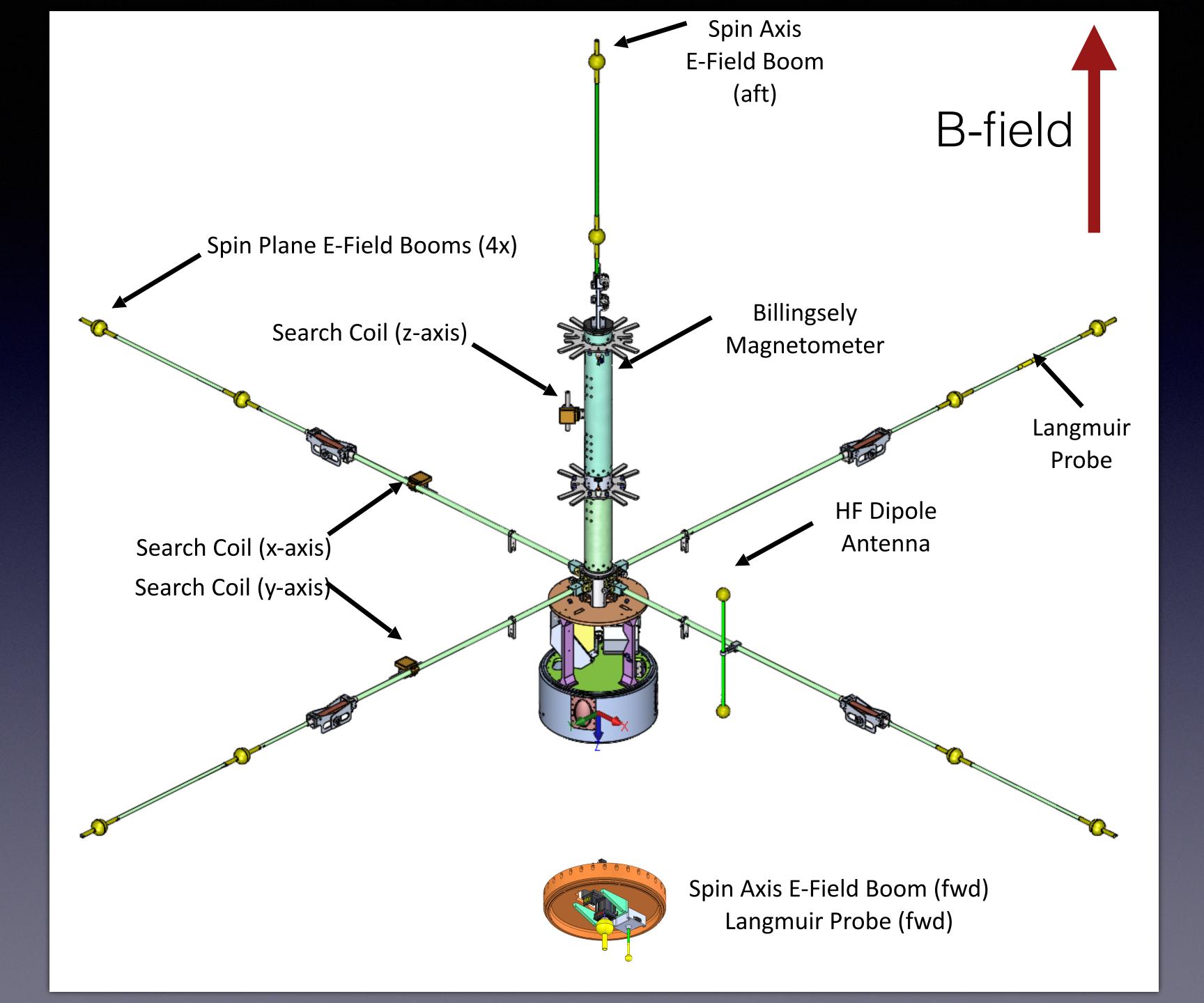


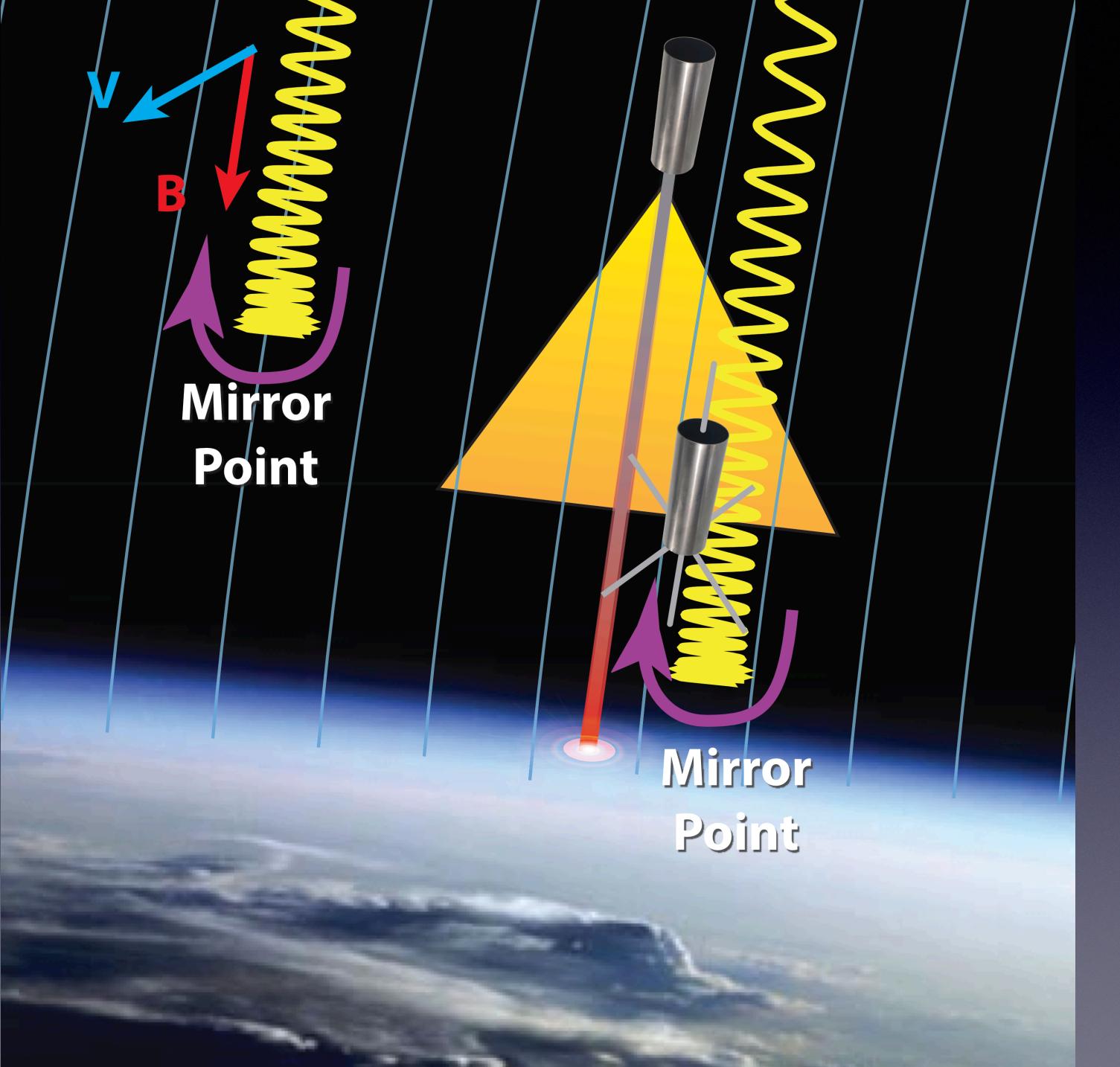




#### GSFC/Pfaff, Farrell

- E-field boomsDual-hinge fiberglassFold-out axial boomStacer boom (axial)
- -- Langmuir Probes
- -- Search Coil sensors
- -- Billingsley Magnetometer
- -- Fields Electronics Box



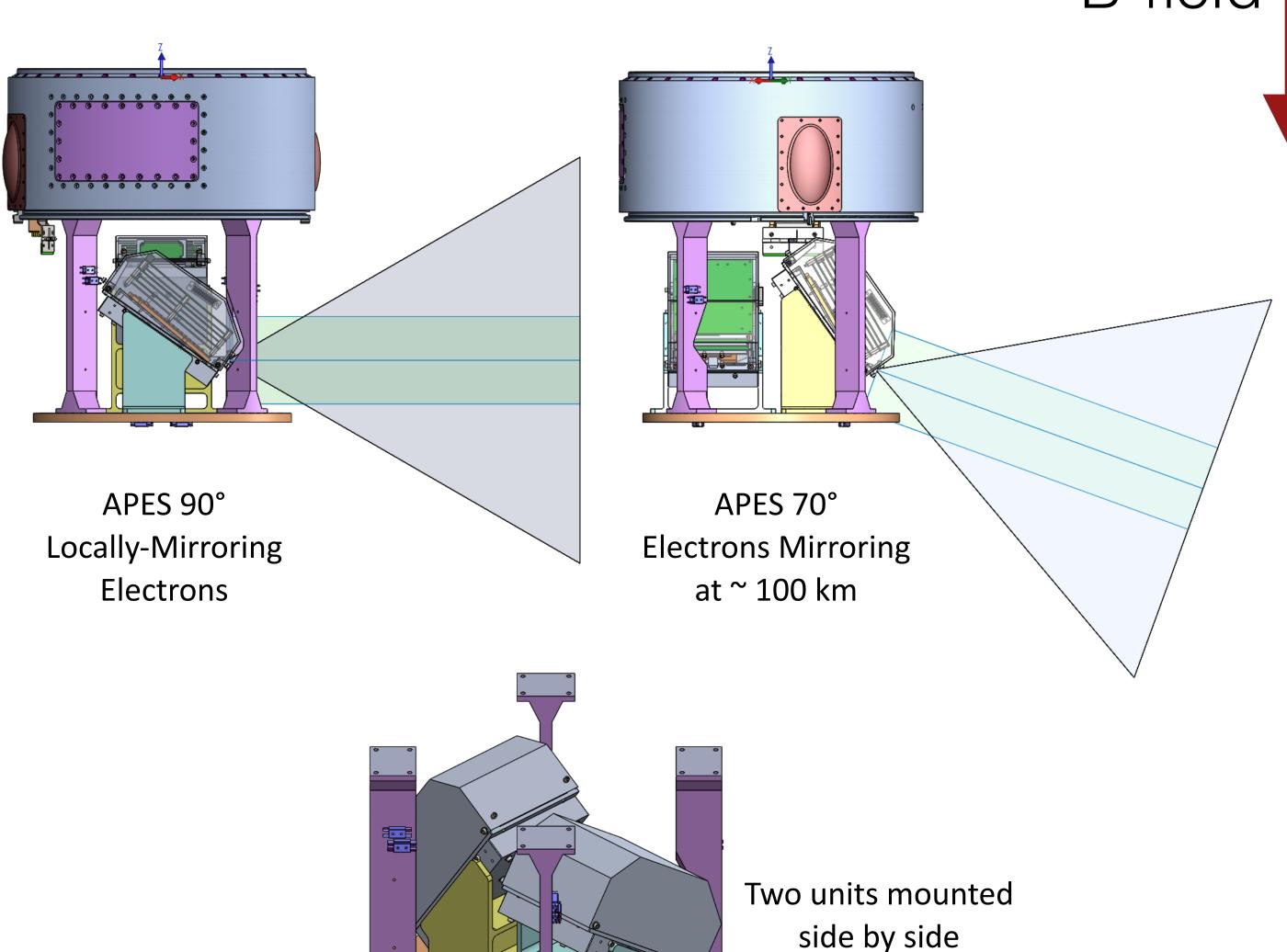


- At 300-500 km most ambient electrons (10s keV) mirror above the receiver
- We look for scattering by comparing the flux of particles mirroring locally (90°) and below the receiver (70°)
- Compare flux ratio with beam ON vs beam OFF

### GSFC/Samara

- -- APES electron detector
- -- HF Analog
  Experiment

#### B-field



# Parameters that determine wave characteristics

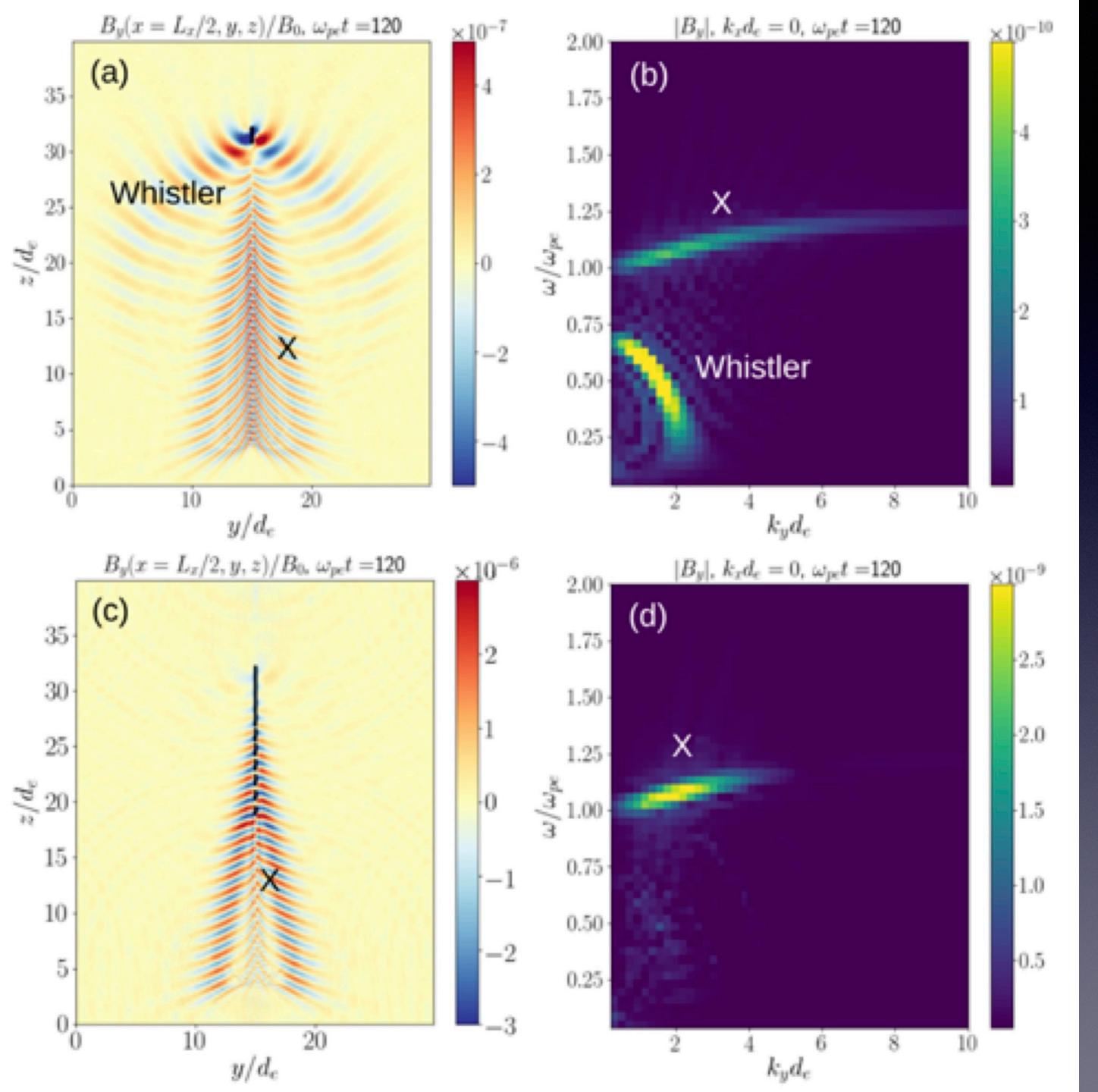
- Beam Energy
- Beam Frequency
- Beam Duty Cycle

Pulse width in microseconds						Period in microseconds	Frequency in Hertz
1	2.5%	2	5.0%	4	10.0%	40	25,000
1	2.0%	2	4.0%	5	10.0%	50	20,000
3	3.0%	5	5.0%	10	10.0%	100	10,000
12	2.4%	25	5.0%	50	10.0%	500	2,000

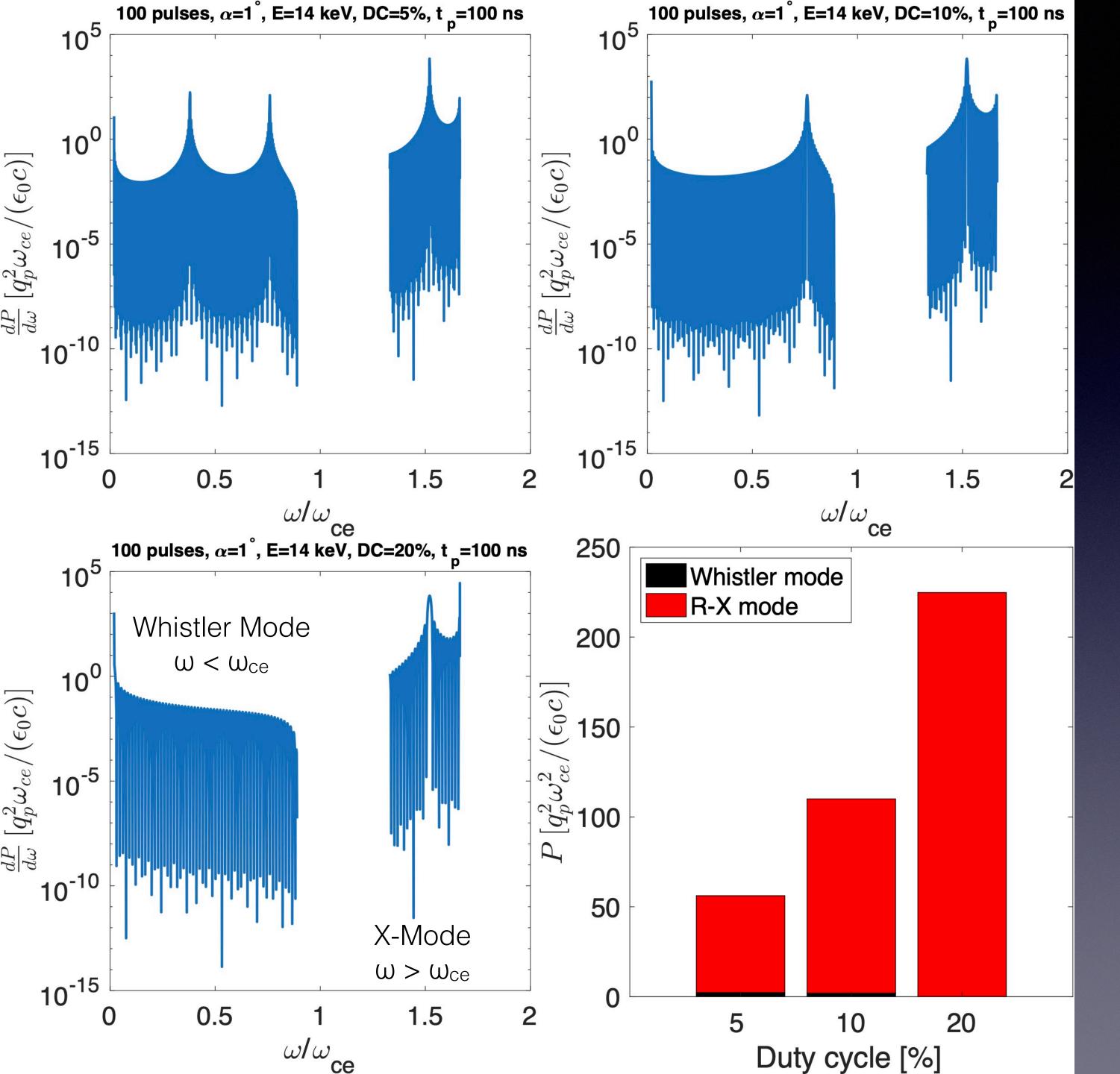
Ambient Plasma conditions (B, density, temperature)

Can repeat all of these every 12 seconds ≈ 25 complete repeated sequences

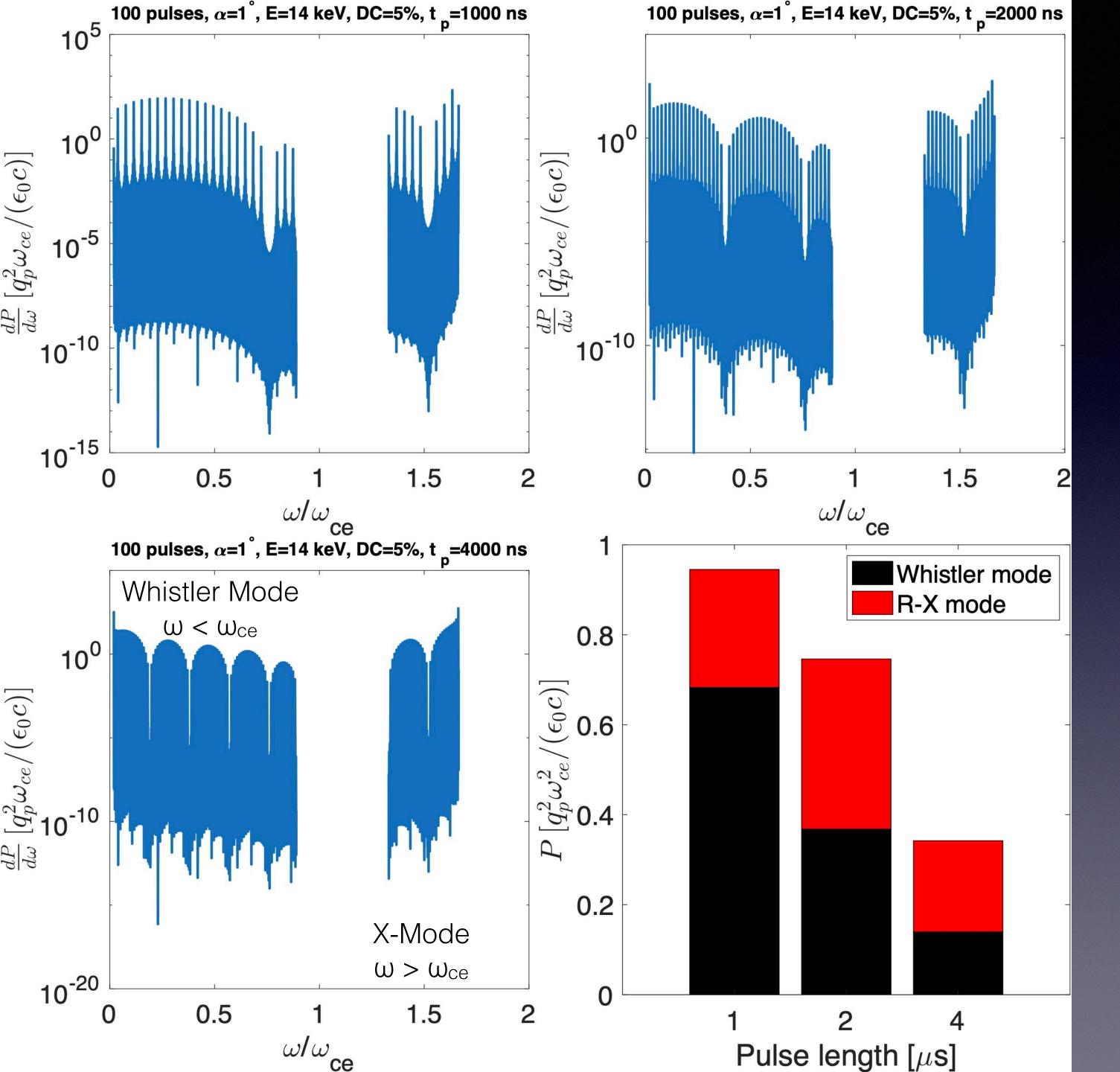
- Prediction of Wave Characteristics as a function of beam pulsing parameters.
- One new tool is Spectral Plasma Solver (SPS) methods.
   [Delzanno et al., 2015]
- SPS is like PIC but solves the equations in the spectral domain
- Well suited for wave-particle interactions because of inherently low numerical noise



- SPS simulations show both X-mode and Whistler-mode waves will be generated
- Different pulsing schemes will determine Whistler & X mode spectra
- Whistler mode waves have a broad wave cone
- X mode waves may only be detected at small cross-field separations



- Dependence on Duty Cycle
- 5% to 20% duty cycle
- 14 keV, 100 ns duration
- Blue curves show power spectrum with harmonics at N times pulsing frequency
- Whistler mode at  $\omega < \omega_{ce}$
- X Mode at  $\omega > \omega_{ce}$



- Dependence on Pulse Duration
- 1 s to 4 s pulse duration
- 14 keV, 5% duty cycle
- Blue curves show power spectrum with harmonics at N times pulsing frequency
- Short pulse durations (higher frequency modulation) produce strong X Mode response
- Minimum duration limited by RF cavity filling time (not yet known)

## "Conclusions"

- If everything works the results should be really cool
- The methodology allows very sensitive and quantitative tests for quasi-linear theory - both for wave generation by particles and particle scattering by waves
- We can and will measure waves generated by non-linear processes
- Narrow-band waves are more likely to produce non-linear pitch angle scattering which will also be tested
- Electron beams are currently being investigated for use in Radiation Belt Remediation [Reeves et al., 2018, 2020]